

Notes for Lecture 12

Gravity

In the previous lecture note, we wrote down Newton's law of gravity in several ways. Here in this note, we explore three examples. The first example is a simple integration problem, while in the other two we make use of the Gauss law.

Here is a question for extra work for motivated students. Consider Earth as a sphere with uniform density. Drill a straight hole through it, passing through the center of the Earth, and drop a mass into that hole. (a) What kind of motion will that mass go through, if we ignore friction. (b) How long will it take to come back to the original drop-off point? As far as figuring out the force field is concerned, this problem is an easier version of Example 12.3, presented below. Now, let us add a twist to this problem. Supposed that the drilled path does *not* go through the center (and so the path will be a shorter one now). Answer questions (a) and (b) for this new straight path, again assuming no friction (however, some normal force *is* required). Let us assume that the Earth is not spinning (although for the first part, the spinning would not matter if the spin axis coincides with the drilled path).

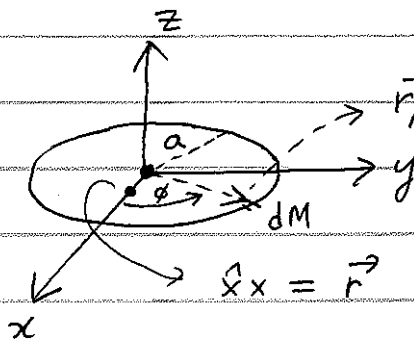
Note for the first example: Here, the problem is considered for \vec{r} in the xy plane. In particular, we set $y = 0$, and evaluate the gravitational potential Φ as a function of x . When done with the calculation, we can replace x with $\rho = \sqrt{x^2 + y^2}$, thanks to the cylindrical symmetry. So, the solution given $\Phi(x, y, z = 0) = -\frac{MG}{2\pi} \int_0^{2\pi} d\phi \frac{1}{\sqrt{a^2 - 2ax \cos \phi + x^2}}$ can be rewritten as

$$\Phi(x, y, z = 0) = -\frac{MG}{2\pi} \int_0^{2\pi} d\phi \frac{1}{\sqrt{a^2 - 2a\rho \cos \phi + \rho^2}}, \quad \rho = \sqrt{x^2 + y^2}. \quad (12.1)$$

For your optional extra work, you can prove that the procedure $x \rightarrow \rho$ is a valid one by a more explicit (but very simple) calculation. Also, it is left for your extra work to write down $\Phi(x, y, z)$ for any z , and prove that the origin is a *saddle point*.

Example 12.1

Gravity example

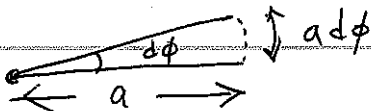


• Consider \vec{r} in the xy plane only.

$$\Phi = - \int dM \frac{G}{|\vec{r} - \vec{r}_M|}$$

$$\frac{dM}{a d\phi} = \frac{M}{2\pi a} \leftarrow \begin{array}{l} \text{total mass} \\ \text{circumference} \end{array} = \text{mass density (linear)}$$

$$dM = M \cdot \frac{d\phi}{2\pi}$$



$$|\vec{r} - \vec{r}_M| = \sqrt{(a \cos \phi - x)^2 + a^2 \sin^2 \phi}$$

$$= \sqrt{a^2 - 2a \cos \phi x + x^2}$$

$$\Phi = - \frac{MG}{2\pi} \int d\phi \frac{1}{\sqrt{a^2 - 2a \cos \phi x + x^2}}$$

If $|x/a| \ll 1$,

$$\frac{1}{\sqrt{a^2 - 2a \cos \phi x + x^2}} \approx \frac{1}{a} \left(1 - 2 \cos \phi \frac{x}{a} + \left(\frac{x}{a}\right)^2 \right)^{-\frac{1}{2}} \quad \frac{1}{2} \cdot \frac{3}{2} \cdot 4$$

$$\approx \frac{1}{a} \left[1 + \frac{1}{4} \cos \phi \frac{x}{a} - \frac{1}{2} \left(\frac{x}{a}\right)^2 + \frac{3}{2} \cos^2 \phi \left(\frac{x}{a}\right)^2 \right]$$

Integrating,

$$\Phi \approx - \frac{MG}{a} \left(1 + \frac{1}{4} \left(\frac{x}{a}\right)^2 \right)$$

One can change $x \rightarrow \rho = \sqrt{x^2 + y^2}$
 considering the ~~symmetry~~ symmetry of the prob.

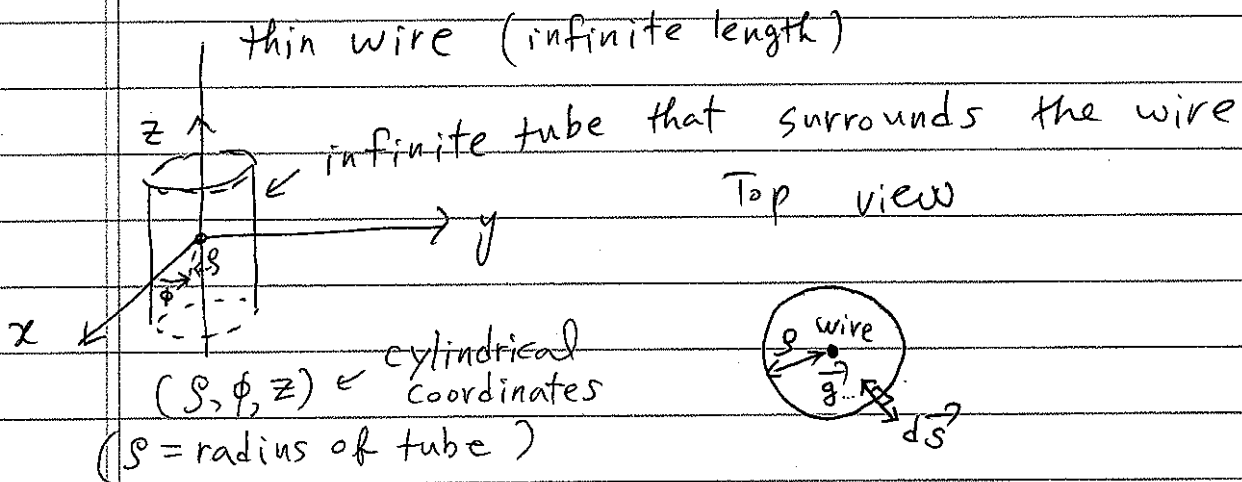
$$\Phi \approx - \frac{MG}{a} \left(1 + \frac{1}{4} \left(\frac{\rho}{a}\right)^2 \right) \leftarrow \begin{array}{l} (z=0) \\ \text{case} \\ \text{only} \end{array}$$

The origin is an unstable equilibrium $\left(\frac{\rho}{a}\right) \ll 1$
 point.

Example 12.2

Gravity example #2

Gauss law



Considering a tube (infinite length) of radius

$$\rho = \sqrt{x^2 + y^2}$$

and using Gauss law for it

$$\int_{\text{tube}} \vec{\nabla} \cdot \vec{g} \, dV = \int_{\text{tube surface}} \vec{g} \cdot d\vec{S} = -g \int_0^{2\pi} \int_0^L \rho \, d\phi \, dz$$

length $L \rightarrow \infty$
of tube or wire

the magnitude of \vec{g}
constant on the tube
due to the cylindrical
symmetry

$$= -g L 2\pi \rho$$

$$= -4\pi G M$$

M : total mass of the wire ($\rightarrow \infty$)

L : length of the wire ($\rightarrow \infty$)

$A \equiv M/L = \text{linear density of the wire}$
 (finite)
 \rightarrow mass per unit length

$$g = 2GA/\rho$$

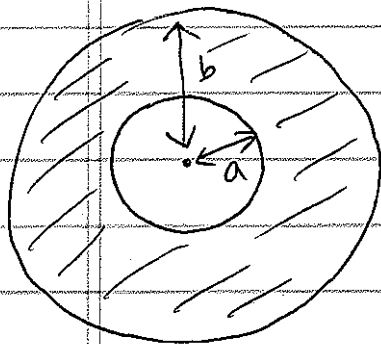
$$\vec{g} = -2GA \hat{\rho}/\rho$$

- sign means
 \vec{g} towards the origin

$$\Phi = +2GA \ln \rho$$

Example 12.3

Spherically symmetric problem.
Gauss law problem!



Total mass M

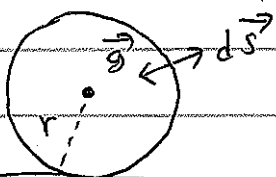
mass between $a < r < b$ only
mass "shell"

Due to the spherical symm.

$g = |\vec{g}|$, Φ depend only
on r
not on θ, ϕ

(r, θ, ϕ) --- spherical coord. sys.

Gauss law at any r



$$-g \cdot 4\pi r^2 = \text{mass inside radius } r \times (-4\pi G)$$

① If $r < a$, $g = 0$ \because no mass inside

② If $r > b$, $g = \frac{GM}{r^2}$ $\vec{g} = -\frac{GM}{r^2} \hat{r}$

Just as though there is a point mass at the origin!

\Rightarrow Generally true for any spherical mass distribution, if the field is measured completely outside the mass distribution.

③ If $a < r < b$, mass inside $r = \frac{r^3 - a^3}{b^3 - a^3} M$

$$\vec{g} = -\frac{GM}{r^2} \left(\frac{r^3 - a^3}{b^3 - a^3} \right) \hat{r}$$

Note that $\vec{g}(r=a) = 0$ --- agrees with ①

$\vec{g}(r=b) = -\frac{GM}{b^2} \hat{r}$ --- agrees with ②

$\therefore \vec{g}$ is continuous everywhere.

