

PH102, 2012W, Lecture Notes: March 7, Wed, Class 21

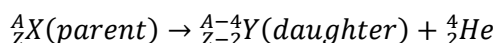
Radioactive decay:

- Definition: An unstable nucleus (parent nucleus) can spontaneously emit small particles or energies to become a nucleus (daughter nucleus) in a more stable state.
- Energy is conserved in radioactive decay: Q (*kinetic energy released*) = $(m_i - m_f)c^2$

Three forms of radioactive decay depending upon what is emitted during radioactive decay

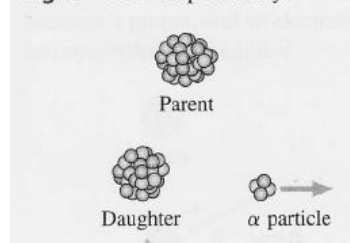
Alpha decay

- Emits an alpha particle (He nucleus=2 protons + 2 neutrons)
- This process makes:



- $Z_{\text{daughter}} = Z_{\text{parent}} - 2$
- $N_{\text{daughter}} = N_{\text{parent}} - 2$
- $A_{\text{daughter}} = A_{\text{parent}} - 4$

Figure 11.18 Alpha decay.



- Released kinetic energy (Q)

$$Q = (m_{\text{parent}} - m_{\text{daughter}} - m_{\frac{1}{2}\text{He}}) c^2$$

- Example: alpha decay of ${}^{238}_{92}\text{U}$

$$\begin{aligned} {}^{238}_{92}\text{U} &\rightarrow {}^{234}_{90}\text{Th} + {}^4_2\text{He} \\ Q &= (m_{\text{parent}} - m_{\text{daughter}} - m_{\frac{1}{2}\text{He}}) c^2 \\ &= (238.050784 - 234.043593 - 4.002603)uc^2 \\ &= 0.004588 \times 931.5 \text{ MeV} = 4.27 \text{ MeV} \end{aligned}$$

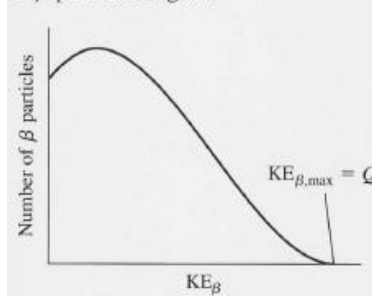
Since the mass of He is a lot smaller than that of Th, the alpha particle gets almost all the kinetic energy during the alpha decay process.

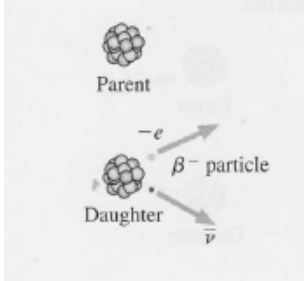
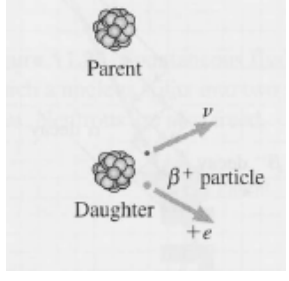
Beta decay emits temporarily created negatively charged electrons (β^- particle) or positively charged electrons (β^+ particle).

In order to satisfy charge conservation, energy conservation, and angular momentum conservation, a new particle is introduced called neutrinos/antineutrinos.

- Charge conservation: since beta particles carry a charge, the charge of the parent nucleus should increase by 1 in β^- decay and decrease by 1 in β^+ decay. As a result, neutrinos' charge should be neutral.
- Angular momentum conservation: since β particles have a spin of $\frac{1}{2}$, to conserve angular momentum, neutrinos should have a spin of $\frac{1}{2}$.
- Energy conservation. Figure 11.19 shows that kinetic energies of β particles emitted during the decay vary greatly from zero to Q (maximum allowed). Therefore, neutrinos should carry portions of kinetic energies with β particles. To allow β particles to have the maximum allowed kinetic energy, neutrinos' mass should be negligible.

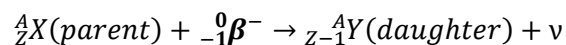
Figure 11.19 The mysterious variation in β particle energies.



β^- decay	β^+ decay
<ul style="list-style-type: none"> An electron (β^- particle) is temporarily created and then emitted. Emits an electron and an anti-neutrino. Changes a neutron inside the nucleus into a proton. 	<ul style="list-style-type: none"> Emits a positron and a neutrino Changes a proton inside the nucleus into a neutron.
${}^A_Z X(\text{parent}) \rightarrow {}^A_{Z+1} Y(\text{daughter}) + {}^0_{-1} \beta^- + \bar{\nu}$ <ul style="list-style-type: none"> $Z_{\text{daughter}} = Z_{\text{parent}} + 1$ $N_{\text{daughter}} = N_{\text{parent}} - 1$ $A_{\text{daughter}} = A_{\text{parent}}$ <p>Released kinetic energy</p> $Q = (m_{\text{parent}} - m_{\text{daughter}}) c^2$	${}^A_Z X(\text{parent}) \rightarrow {}^A_{Z-1} Y(\text{daughter}) + {}^0_{+1} \beta^+ + \nu$ <ul style="list-style-type: none"> $Z_{\text{daughter}} = Z_{\text{parent}} - 1$ $N_{\text{daughter}} = N_{\text{parent}} + 1$ $A_{\text{daughter}} = A_{\text{parent}}$ <p>Released kinetic energy</p> $Q = (m_{\text{parent}} - m_{\text{daughter}} - 2m_{\text{electron}}) c^2$
<p>Example:</p> ${}^{12}_5 B \rightarrow {}^{12}_6 C + {}^0_{-1} \beta^- + \bar{\nu}$ $Q = (12.014352 - 12)uc^2 = 13.4 \text{ MeV}$	${}^{12}_7 N \rightarrow {}^{12}_6 C + {}^0_{+1} \beta^+ + \nu$ $Q = (12.018613 - 12 - 2 \times 0.0005486)uc^2 = 16.3 \text{ MeV}$
	

The third form of beta decay is Electron Capture.

- A nucleus with too many protons can change a proton into a neutron by capturing an electron.
- Electron capture is easier than β^+ decay since an electron is already exists for a nucleus to capture.

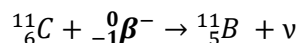


- $Z_{\text{daughter}} = Z_{\text{parent}} - 1$
- $N_{\text{daughter}} = N_{\text{parent}} + 1$
- $A_{\text{daughter}} = A_{\text{parent}}$

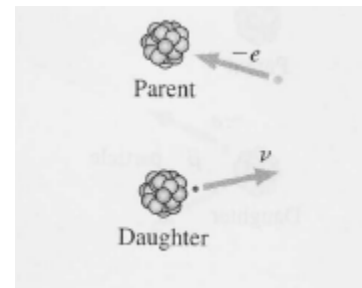
- Released kinetic energy

$$Q = (m_{\text{parent}} - m_{\text{daughter}}) c^2$$

- Example:



$$Q = (11.01143 - 11.009305)uc^2 = 1.97 \text{ MeV}$$



Gamma Decay

- A nucleus in an excited state emits photons (gamma particles, γ) to go into a lower energy state.
- Gamma decay does not alter N or Z.
- Gamma energies are characteristic of a given isotope, and are thus used to identify the isotope.

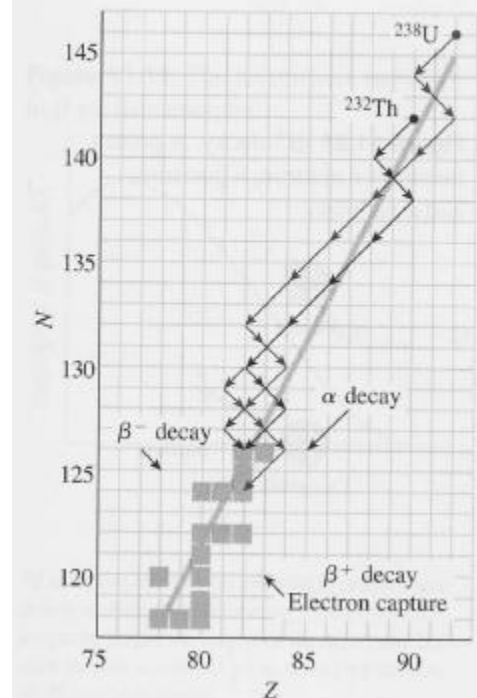
Figure 11.24 Gamma decay.



Decay series

- An unstable nucleus can be involved in a series of decays until it finds a stable state.
- We can plot this process on a graph that represents N and Z numbers of each nucleus in the series.
- Figure 11.23 shows such a graph. The gray line represents the line of stability. The figure shows two series: one for U-238 and the other for Th-232.
- Alpha decay is shown by an arrow
 - $Z_{daughter} = Z_{parent} - 2$
 - $N_{daughter} = N_{parent} - 2$
- β^- decay
 - $Z_{daughter} = Z_{parent} + 1$
 - $N_{daughter} = N_{parent} - 1$
- β^+ decay and electron capture
 - $Z_{daughter} = Z_{parent} - 1$
 - $N_{daughter} = N_{parent} + 1$

Figure 11.23 The “directions” of α and β decays, and the decay series of uranium-238 and thorium-232.



Radioactive Decay Law

- For all decays, the rate of decay over time will be proportional to the sample size:

$$\frac{dN}{dt} \propto N$$

$$\frac{dN}{dt} = -\lambda N \quad \text{where } N = \text{Number of nuclei; } \lambda = \text{decay constant}$$

$$\frac{dN}{N} = -\lambda dt$$

$$\int_{N_0}^N \frac{dN}{N} = -\lambda \int_0^t dt$$

$$\ln \frac{N}{N_0} = -\lambda t$$

$$N = N_0 e^{-\lambda t}$$

- Decay rate $R = \lambda N$ (decays per second)
- A sample of the same nuclei will decay by the same fraction in equal successive intervals of time.
- Half-life ($T_{1/2}$) is defined as the time interval at which half of the sample will decay:

$$\frac{1}{2} N_0 = N_0 e^{-\lambda T_{1/2}}$$

From this relationship, we can calculate decay constant

$$\lambda = \frac{\ln 2}{T_{1/2}}$$

Figure 11.26 shows the relationship between half life and the number of nuclei remaining over time.

- Half-lives vary widely, from 10^{-22} seconds to 10^{+17} years.

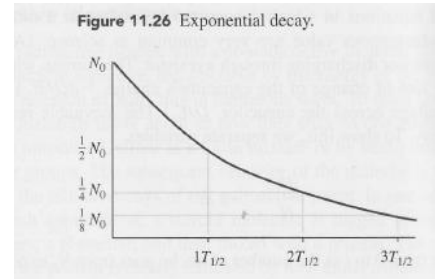


TABLE 11.3 Selected decays

Isotope	Decay Mode	Half-Life
$^{35}_{20}\text{Ca}$	β^+	50 ms
^3_1H	β^-	12.3 yr
$^{238}_{92}\text{U}$	α	4.5×10^9 yr

Example:

A sample holds 2 μg of tritium.

- Initial decay rate (R)

$$R = \lambda N = \frac{\ln 2}{T_{1/2}} \cdot \frac{\text{sample mass}}{\text{atomic mass of Tritium}}$$

Since Tritium's atomic mass = 3.02 u and 1 u = $1.66 \times 10^{-27} \text{ kg}$, $T_{1/2} = 12.3 \text{ yr} = 12.3 \times 3.16 \times 10^7 \text{ sec}$

$$\lambda = \frac{\ln 2}{T_{1/2}} = 1.78 \times 10^{-9} \text{ /sec}$$

$$R = 7.1 \times 10^8 \text{ decays/sec}$$

- Elapsed time before the decay rate falls to 1% of its initial value?

$$N = N_0 e^{-\lambda t}$$

$$\frac{1}{100} N_0 = N_0 e^{-\lambda t}$$

$$t = \frac{-\ln(\frac{1}{100})}{\lambda} = 2.6 \times 10^9 \text{ sec} = 81.7 \text{ years}$$

Carbon-14 dating

- Carbon-14's β^- decay has a half-life of 5730 years.
- Carbon-14 dating only works for formerly living organisms.
- Carbon-14's amount is constantly maintained for living organisms since living organisms exchange Carbon with the environment. Ratio of naturally produced C14/C12 = 1.3×10^{-12}
- When, a living organism dies, it stops the exchange process, thus C-14 in the dead organism decays exponentially.

Example:

What is the age of a fossil sample that contains 6 g of carbon and has a decay rate (R) of 30 decays per minute?

- The sample has 6 g of Carbon. 1.3×10^{-12} th of 6g of Carbon should be C-14 at the time when the organism in the sample was alive. This amount of C-14 is subject to the decay. Therefore,

$$N_0(C_{14}) = (1.3 \times 10^{-12}) \cdot \left(\frac{1}{2}\right) \cdot (6.02 \times 10^{23}) = 3.9 \times 10^{11}$$

- Decay constant

$$\lambda = \frac{\ln 2}{T_{1/2}} = \frac{\ln 2}{5730 \text{ years} \cdot 3.16 \times 10^7 \text{ sec/year}} = 3.83 \times 10^{-12} \text{ /sec}$$

- Current decay rate (R)=30 decays/60 seconds=1/2 decays/sec

$$R = N \cdot \lambda = N \cdot 3.83 \times 10^{-12} \text{ /sec}$$

$$N = 1.31 \times 10^{11}$$

$$N = N_0 e^{-\lambda t}$$

$$t = -\frac{1}{\lambda} \ln \frac{N}{N_0} = -\frac{1}{3.83 \times 10^{-12} \text{ /sec}} \ln \frac{1.31 \times 10^{11}}{3.9 \times 10^{11}} = 2.86 \times 10^{11} \text{ sec} \sim 9000 \text{ years}$$